# ON THE RELATIVE PERIODIC MOIIONS OF A PENDULUM 

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Considered are the properties of the relative periodic motions of a rigid body suspended on an elastic string in a uniformly rotating vertical plant. The relative periodic motions of a mathematical pendulum with an elastic :tring were considered in [1] and [2].

1. Let us denote by $0 x y$ a system of coordinates rotating uniformly with respect to the downward oy-axis, relative to which the motion of the rigid body will be studied. The elastic string, considered as a linear ma: ilees spring with an elastic constant $c$, iu attached at a point $O_{1}$ (See Fig.1) where $\delta O_{1}=h$. The angle of deflection of the strine from the vertical axis $0_{1} \mu_{1}$ will be denoted by $r_{1}$ anu ifi icneth by $p$. A body with a mass $m$ is suspenued from the spring at point $0_{2}$. The angle between the straight line passing through $O_{2}$ and the center of gravity $C$ of the body and the vertical will be denoted as $\varphi_{2}$, and let us assume $O_{2} C=a$. Choosing a body system of coordinates $C \bar{\Pi} \Pi$, such that $C \eta$ passes through $O_{2}, C \zeta$ is orthogonal to $C \eta$ and lying in the plane $O x y$, while $C \zeta$ 1s orthogonal to oxy.


We assume that the axes of the system $0 \varepsilon \eta \delta$ are the principal axes of inertia of the body. The principal moments of inertia with respect to the axes $C \xi, C \eta$ and $C S$ will be denoted by $J_{v}(v=1,2,3)$, respectively.

The kinetic energy of the system in its absulute motion. according to the Koenig theorem, is

$$
T=1 / 2 m c_{r}^{2}+1 / 2\left|J_{3} \varphi_{2}^{2}+\omega^{2}\left(J_{1} \sin ^{2} \varphi_{2}+J_{2} \cos ^{2} \varphi_{2}\right)\right|
$$

where $v$ is the abroll, whocicy of point $C$. In view "fl the fact that the coordinates of the center of gravity of the body relative to the $0 x y$ system are defined by Expressions

$$
\begin{gathered}
x_{c}=h+\rho \sin \varphi_{1}+a \sin \varphi_{2} \\
y_{c}=\rho \cos \varphi_{1}=a \cos \varphi_{2}
\end{gathered}
$$

and that

$$
\tau_{c}^{2}=x_{r}^{2}+y_{c}^{2}+\omega^{2} x_{c}^{2}
$$

the erpresslon for the kinetic energy becomes

$$
\begin{gather*}
T=1 / 2 m\left[\rho^{2}+\rho^{2} \varphi_{1}^{2}+a \varphi_{2}^{2}+2 a \rho \varphi_{2} \sin \left(\varphi_{1}-\varphi_{2}\right)+2 a \rho \varphi_{1} \varphi_{2} \cos \left(\varphi_{1}-\varphi_{2}\right)+\right. \\
\left.+\omega^{2}\left(h+\rho \sin \varphi_{1}+a \sin \varphi_{2}\right)^{2}\right]+1 / 2\left[J_{3} \varphi_{2}^{2}+\omega^{2}\left(J_{1} \sin ^{2} \varphi_{2}+J_{2} \cos ^{2} \varphi_{2}\right)\right] \tag{1.1}
\end{gather*}
$$

The putential energy of the system i.e. the sum of the energy of elastic deformation of the itrime and the gravitational energy, is given by the Expmain

$$
\begin{equation*}
\Pi=1 / 2 c(\rho-l)^{2}-m g\left(\rho \cos \varphi_{1}+a \cos \varphi_{2}\right) \tag{1.2}
\end{equation*}
$$

where $i$ is the free length of the string.
Beariff in mind (1.1) and (1.2) we obtain, according to the Lagrange quationt , the system of differential equations of motion

$$
\rho^{*}+\rho\left(k^{2}-\varphi_{1}^{2}-\omega^{2} \sin ^{2} \varphi_{1}\right)+a \varphi_{2}^{*} \sin \left(\varphi_{1}-\varphi_{2}\right)-a \varphi_{2}{ }^{* 2} \cos \left(\varphi_{1}-\varphi_{2}\right)-
$$

$$
-\omega^{2}\left(h+a \sin \varphi_{2}\right) \sin \varphi_{1}-g \cos \varphi_{1}-k^{2} l=0
$$

$$
\rho \varphi_{1}^{\prime \cdot}+2 \rho^{\prime} \varphi_{1}^{\cdot}+a \varphi_{2}^{\prime} \cos \left(\varphi_{1}-\varphi_{2}\right)+a \varphi_{2} \sin \left(\varphi_{1}-\varphi_{2}\right)-\omega^{2}\left(h+\rho \sin \varphi_{1}+\right.
$$

$$
\begin{equation*}
\left.+a \sin \varphi_{2}\right) \cos \varphi_{1}+g \sin \varphi_{1}=0 \tag{1.3}
\end{equation*}
$$

$l_{1} \varphi_{2}{ }^{*}+\left(\rho \varphi_{1}{ }^{*}+2 \rho \varphi_{1}\right) \cos \left(\varphi_{1}-\varphi_{2}\right)-\left(\rho^{*}-\rho \varphi_{1}{ }^{2}\right) \sin \left(\varphi_{1}-\varphi_{2}\right)-$
$-\omega^{2}\left(h+\rho \sin \varphi_{1}+a \sin \varphi_{2}\right) \cos \varphi_{2}+\left[\left(J_{2}-J_{1}\right) / 2 a m \mid \sin 2 \varphi_{2}+g \sin \varphi_{2}=0\right.$
where

$$
l_{1}=\frac{1}{a}\left(a^{2}+\frac{J_{3}}{m}\right), \quad k^{2}=\frac{c}{m}
$$

Here 12 is the derived length of the physical pendulum (body) relative to the puint $O_{2}$.

Let us consider the oscillations of the system near the position of relative equilibrium. Let us assume

$$
\begin{equation*}
J_{1}=J_{2} \tag{1.4}
\end{equation*}
$$

It can be proved that if (1.4) is valid, then for the state of relative equilibrium the angles $p_{10}$ and $\varphi_{\infty}$ are equal to each other and, consequently, the following substitution can be made:

$$
\begin{equation*}
p(t) \therefore b+\xi(t), \quad \varphi_{1}(t)=\varphi_{0}+\varphi(t), \quad \varphi_{2}(t)=\varphi_{0}+\psi(t) \tag{1.5}
\end{equation*}
$$

Here mo is the value of the angles $m_{10}$ and $q_{20}$, and $k$ is the length of the pendulum string in relative equilibrium. The quantities fo and $b$ are determined by the equalities

$$
\begin{gather*}
k^{2}(b-l)=\omega^{2}\left(h+b \sin \varphi_{0}+a \sin \varphi_{0}\right) \sin \varphi_{0}+g \cos \varphi_{0}  \tag{1.6}\\
g \sin \varphi_{0}=\omega^{2}\left(h+b \sin \varphi_{0}+a \sin \varphi_{0}\right) \cos \varphi_{0}
\end{gather*}
$$

Bearins in mind (1.4) and (1.6), by the substitution of (1.5) into (1.3), we obtain

$$
\begin{equation*}
\xi^{\prime \prime}+\left(k^{2}-\omega^{2} \sin \varphi_{0}\right) \xi-1 / \omega^{2} \omega^{2} b \sin 2 \varphi_{0} \varphi-1 / 2 \omega^{2} a \sin 2 \varphi_{\psi} \psi=f_{1}+\ldots \tag{1.7}
\end{equation*}
$$

$b \varphi^{*}+a \varphi^{\prime \prime}-\omega^{2} a \cos ^{2} \varphi_{0} \xi+\left[k^{2}(b-l)-\omega^{2} b \cos ^{2} \varphi_{0} I \varphi-\omega^{2} a \cos ^{2} \varphi_{0} \psi=\dot{f}_{2}+\ldots\right.$
$l_{1} \psi^{\prime \prime}+b \varphi^{\prime \prime}-1 / 2 \omega^{2} \sin 2 \varphi_{0} \xi-\omega^{2} b \cos ^{2} \varphi_{0} \varphi+\left[k^{2}(b-l)-\omega^{2} a \cos ^{2} \varphi_{0}\right] \psi=f_{3}+\ldots$ where

$$
\begin{gather*}
f_{1}=b \varphi^{2}-a(\varphi-\psi) \psi^{\prime \prime}+\omega^{2} \sin 2 \varphi_{0} \xi \varphi+\omega^{2} a \cos ^{2} \varphi_{0} \varphi \psi-1 / 2 \omega^{2} a \sin ^{2} \varphi_{0} \psi^{2}+a \varphi^{2}+ \\
+1 / 2\left(2 \omega^{2} b \cos 2 \varphi_{0}-a \omega^{2} \sin ^{2} \varphi_{0}-\omega^{2} h \sin \varphi_{0}-g \cos \varphi_{0}\right) \varphi^{2} \\
f_{2}=-\xi \varphi^{\circ}-2 \xi \varphi+\omega^{2} \cos 2 \varphi_{0} \xi \varphi-1 / \omega^{2} a \sin 2 \varphi_{0} \varphi \varphi-3 / 4 \omega^{2} b \sin 2 \varphi_{0} \varphi^{2}-1 / 2 \omega^{2} a \sin 2 \varphi_{0} \psi^{2} \\
f_{3}=-\xi \varphi^{\prime \prime}-(\varphi-\psi) \xi-2 \xi \varphi^{\circ}+\omega^{2} \cos ^{2} \varphi_{0} \xi \varphi-1 / 2 \omega^{2} b \sin 2 \varphi_{0} \varphi \psi- \\
-\omega^{2} \sin ^{2} \varphi_{0} \xi \psi-1 / 4 \omega^{2} b \sin 2 \varphi_{0} \varphi^{2}-3 / 4 \omega^{2} \sin 2 \varphi_{0} \psi^{2} \tag{1.8}
\end{gather*}
$$

Neglecting the nonlinear terms we obtain the eyntum

$$
\begin{array}{r}
\xi^{*}+a_{11} \xi+a_{12} \varphi+a_{13} \psi=0 \\
b \varphi+\varphi^{\prime \prime}+a^{\prime *}+b_{11} \xi+b_{12} \varphi+b_{13} \psi=0  \tag{1.9}\\
l_{1} \psi^{*}+b \varphi^{*}+c_{11} \xi+c_{12} \varphi+c_{13} \psi=0
\end{array}
$$

where

$$
\begin{gather*}
a_{11}=k^{2}-\omega^{2} \sin ^{2} \varphi_{0}, \quad b_{11}=-\omega^{2} a \cos ^{2} \varphi_{0}, \quad c_{11}--1 / 2 \omega^{2} \sin 2 \varphi_{0} \\
{ }^{2}=-1 / 2 \omega^{2} b \sin 2 \varphi_{0}, \quad b_{12}=k^{2}(b-l)-\omega^{2} b \cos ^{2} \varphi_{0}, \quad c_{12}=-\omega^{2} b \cos ^{2} \varphi_{0} \\
a_{13}=-1 / 2 \omega^{2} a \sin 2 \varphi_{0}, \quad b_{13} \cdots-\omega^{2} a \cos ^{2} \varphi_{0}, \quad c_{13}=h^{2}(b-l)-\omega^{2} a \cos ^{2} \varphi_{0} \tag{1.10}
\end{gather*}
$$

The fundamental equation of the system will be

$$
\left|\begin{array}{ccc}
a_{11}+r^{2} & a_{12} & a_{13}  \tag{1.11}\\
b_{11} & b_{12}+b r^{2} & b_{13}+a r^{2} \\
c_{11} & c_{12}+b r^{2} & c_{13}+l_{1} r^{2}
\end{array}\right|=0
$$

The condition that Equation (1.11) would have purely inarsinary foul: l:: reduced to the inequality

$$
\begin{align*}
k^{2}\left[g \cdot \cos \varphi_{0}\right. & \left.+\omega^{2}\left(h \sin \varphi_{0}+l \sin ^{2} \varphi_{0}-b \cos ^{2} \varphi_{0}-a \cos 2 \varphi_{0}\right)\right]+ \\
& +\omega^{4} b \cos ^{2} \varphi_{0} \sin \varphi_{0}\left(\sin \varphi_{0}-a \cos \varphi_{0}\right)>0 \tag{1.12}
\end{align*}
$$

Under condition (1.12) the system has three pair: of purcly imatinary roots. On the basis of the general theory of linear equations with contant coefficients, the system ( 1.7 ) can be transformed into a form similar to th. system [3] (p.435)

$$
\frac{d x}{d t}=-r y+X, \quad \frac{d y}{d t}=r x+Y
$$

$$
\frac{d x_{s}}{d t}=b_{81} x_{1}+\ldots+b_{s m^{2}} x_{i n}+a_{s} x+b_{s} y+X_{s} \quad(s=1, \ldots, m ; m=4)
$$

where $\pm$ ir is any pair of purely imaginary rooto.
The propused :yytem may be considered as a Liapunov :yyiem relative tu such a pair of purety imacinary roots. On the baci.. of the theurem [3] ( $p .442$ ), it may be aseerted that the yystem permit: a perivicicolution dependent on an arbitrary parameter. This parameter i: the initial value $z$ of the quantity $r$.

The "basic" paii w fur ly imaginary roote can be any fair and there fore this system permile: thace perivale solutione.
2. Now ict ue utudy the properties of periodic notione near the poritiun of relative equilibrium which have an approximate poriori $z / r_{1}$.

Let the periou of the sulution be of the form

$$
\begin{equation*}
T_{1}=\frac{2 \pi}{r_{0}}\left(1+\delta_{1} \lambda \div \delta_{2} \lambda^{2}+\ldots\right) \tag{2.1}
\end{equation*}
$$

where $\delta_{1}, \delta_{2}, \ldots$ are constants subivet to detcomination and $\alpha=a y$.
We introduce the variable $T$ in place or the variable + ini the cquations by means of the substitution

$$
\begin{equation*}
t=\tau\left(1+\delta_{1} \lambda+\delta_{2} \lambda^{2}+\ldots\right) \tag{2.2}
\end{equation*}
$$

Then the problem is reduced to finding the puriodic : Luilons with the period $z_{1}$ of the syitem

$$
\begin{gather*}
\xi \cdot \cdots+a_{11} \xi+a_{12} \varphi+a_{13} \psi=f_{1}^{*}+\cdots \\
b \varphi^{*}+a \psi{ }^{*}+b_{11} \xi+b_{12} \varphi+b_{13} \psi=f_{2}{ }^{*}+\cdots  \tag{2.3}\\
l_{1} \psi^{\prime \prime}+b \varphi^{*}+c_{11} \xi+c_{12} \varphi+c_{13} \psi=f_{3}{ }^{*}+\ldots
\end{gather*}
$$

whore $f_{1}^{*}, f^{*}{ }^{*}$ and $f_{s}^{*}$ are obtained from (1.8). The derivatives here and in the folluwing are with respect to $\tau$.

The colutions of the system (2.3) are analytic with respect to $\lambda$, and therefore they will be sought in the form of the series

$$
\begin{align*}
& \xi(\tau)=\lambda \xi_{1}(\tau)+\lambda^{2} \xi_{2}(\tau)+\lambda^{3} \xi_{3}(\tau) \& \ldots \\
& \varphi(\tau)=\lambda \varphi_{1}(\tau)+\lambda^{2} \varphi_{2}(\tau)+\lambda^{3} \varphi_{3}(\tau)+\ldots  \tag{2.4}\\
& \Psi(\tau)=\lambda \psi_{1}(\tau)+\lambda^{2} \psi_{2}(\tau)+\lambda^{3} \psi_{3}(\tau) \notin \ldots
\end{align*}
$$

Where $\xi_{v}, \varphi_{\nu}, \psi_{\nu}(v=1,2,3)$ are periodic functions of $\tau$ of the period $2 \pi / r_{1}$ (here $r_{1}$ and $\psi_{2}$ are periodic functions of $T$ and have not the same meaning as in Section 1) satisfying the initial conditions
$\xi_{1}(0)=1, \quad \varphi_{1}(0)=k_{1}, \quad \Psi_{1}(0)=k_{2}, \quad \xi_{v} \cdot(0)=\varphi_{v} \cdot(0)=\varphi_{v} \cdot(0)=0 \quad(v=1,2,3)$
Substituting (2.4) into (2.3) and equating the coefflcients of like powers of $\lambda$ we obtain a system of equations for determination of $\xi_{v}, \varphi_{v}$ and $\psi_{v}$ $(v=1,2,3)$.

For the functions $\xi_{1}, \varphi_{1}$ and $\psi_{1}$ there results a basic system (1.9) which has the obvious solution

$$
\begin{equation*}
\xi_{1}(\tau)=\cos r_{1} \tau, \quad \varphi_{1}(\tau)=k_{1} \cos r_{1} \tau, \quad \psi_{1}(\tau)=k_{2} \cos r_{1} \tau \tag{2.6}
\end{equation*}
$$

where

$$
k_{1}=\frac{\left(a_{11}-\dot{r}_{1}^{2}\right)\left(b_{13}-a r_{1}^{2}\right)-a_{13} c_{11}}{a_{13}\left(b_{12}-b r_{1}^{2}\right)-a_{12}\left(b_{13}-a r_{1}^{2}\right)}, \quad k_{2}=\frac{\left(a_{11}-r_{1}^{2}\right)\left(b_{12}-b r_{1}^{2}\right)-a_{12} c_{11}}{a_{13}\left(b_{12}-b r_{1}^{2}\right)-a_{12}\left(b_{13}-a r_{1}^{2}\right)}
$$

Furthermore, we have

$$
\begin{gather*}
\xi_{2}{ }^{*}+a_{11} \xi_{2}+a_{12} \varphi_{2}+a_{13} \psi_{2}=F_{0}-2 \delta_{1} F_{1} \cos r_{1} \tau+F_{2} \cos 2 r_{1} \tau \\
l \varphi_{2}{ }^{*}+a \psi_{2}{ }^{*}+b_{11} \xi_{2}+b_{12} \varphi_{2}+b_{13} \psi_{2}=C_{0}-2 \delta_{1} G_{1} \cos r_{1} \tau+G_{2} \cos 2 r_{1} \tau  \tag{2.7}\\
l_{1} \psi_{2} \cdot+b \varphi_{2}{ }^{*}+c_{11} \xi_{2}+c_{12} \varphi_{2}+c_{13} \psi_{2}=H_{0}-2 \delta_{1} H_{1} \cos r_{1} \tau+H_{2} \cos 2 r_{1} \tau
\end{gather*}
$$

where the expressions for $F_{v}, G_{\nu}$ and $H_{v}$ are not derived because of their complexity.

The periodic solution (with the period $2 \pi / r_{1}$ ) of the homogeneous system corresponding to (2.7) will be

$$
\begin{gather*}
\xi_{21}=C_{1} \cos r_{1} \tau+D_{1} \sin r_{1} \tau, \quad \varphi_{21}=k_{1}\left(C_{1} \cos r_{1} \tau+D_{1} \sin r_{1} \tau\right) \\
\psi_{21}=k_{2}\left(C_{1} \cos r_{1} \tau+D_{1} \sin r_{1} \tau\right) \tag{2.8}
\end{gather*}
$$

where the unknown constants $C_{1}$ and $D_{1}$ will be determined later.
The particular solution of the system (2.7) is sought in the form

$$
\begin{gather*}
\xi_{22}=: P_{0}+P_{1} \cos r_{1} \tau+P_{2} \cos 2 r_{1} \tau, \quad \varphi_{22}=Q_{0}+Q_{1} \cos r_{1} \tau+Q_{2} \cos 2 r_{1} \tau \\
\psi_{22}=R_{1}+R_{1} \cos r_{1} \tau+R_{2} \cos 2 r_{1} \tau \tag{2.9}
\end{gather*}
$$

Subitituting (2.9) into (2.7) we obtain for $F_{v}, Q_{v}$ and $A_{v}(\nu=0,1,2)$ the system

$$
\begin{align*}
& a_{11} P_{0}+a_{12} Q_{0}+a_{13} R_{0}=F_{0} \\
& b_{11} P_{0}+b_{12} Q_{0}+b_{13} R_{0}=G_{0}  \tag{2.10}\\
& c_{11} P_{0}+c_{12} Q_{0}+c_{13} R_{0}=H_{0}
\end{align*}
$$

$$
\begin{gather*}
\left(a_{11}-r_{1}^{2}\right) P_{1}+a_{12} Q_{1}+a_{13} R_{1}=-2 \delta_{1} F_{1} \\
b_{11} P_{1}+\left(b_{12}-b r_{1}^{2}\right) Q_{1}+\left(b_{13}-a r_{1}^{2}\right) R_{1}=-2 \delta_{1} G_{1}  \tag{2.11}\\
c_{11} P_{1}+\left(c_{12}-b r_{1}^{2}\right) Q_{1}+\left(c_{13}-l_{1} r_{1}^{2}\right) R_{1}=-2 \delta_{1} H_{1} \\
\left(a_{11}-4 r_{1}^{2}\right) P_{1}+a_{12} Q_{2}+a_{13} R_{2}=F_{2} \\
b_{11} P_{2} \downarrow\left(b_{12}-4 b r_{1}^{2}\right) Q_{2}+\left(b_{13}-4 a r_{1}^{2}\right) R_{2}=G_{2}  \tag{2.12}\\
c_{11} P_{2}+\left(c_{12}-4 b r_{1}^{2}\right) Q_{2}+\left(c_{13}-4 l_{1} r_{1}^{2}\right) R_{2}=H_{2}
\end{gather*}
$$

Thus, the problem is reduced to the determination of cunditions for which these systems can be solved with respect to $F_{v}, Q_{v}$ and $R_{v}$.

Since 0 and $\pm 2 i r_{1}$ are not the ruots of Equations (1.11), the system. (2.10) and (2.12) have a unique solution for $P_{0}, Q_{0}, F_{0}$ and $F_{2}, Q_{2}, H_{\text {: }}$, retipectively.

The characteristic equation of the veterminant for the system (2.11) coincides uxactly with Equation (1.11) for which $r= \pm$ ir is a root. It i.. ubvious now that for $\delta_{1}=0$ the system (2.11) has a solution $Q_{1}=k_{1} P_{1}$ and $\boldsymbol{R}_{1}=k_{2} P_{1}$. Thus,

$$
\begin{gather*}
\xi_{2}=P_{0}+P_{1}^{\prime} \cos r_{1} \tau+D_{1} \sin r_{1} \tau+P_{2} \cos 2 r_{1} \tau \\
\varphi_{2}=Q_{0}+Q_{1}^{\prime} \cos r_{1} \tau+k_{1} D_{1} \sin r_{1} \tau+Q_{2} \cos 2 r_{1} \tau  \tag{2.13}\\
\psi_{2}=R_{0}+R_{1}^{\prime} \cos r_{1} \tau+k_{2} D_{1} \sin r_{1} \tau+R_{2} \cos 2 r_{1} \tau
\end{gather*}
$$

where

$$
P_{1}^{\prime}=P_{1}+C_{1}, \quad Q_{1}^{\prime}=Q_{1}+k_{1} C_{1}, \quad R_{1}^{\prime}=R_{1}+k_{2} C_{1}
$$

From condition (2.5) it follows directly that $D_{1}=0$, therefore we can write

$$
\begin{gather*}
\xi_{2}=P_{0}+P_{1}^{\prime} \cos r_{1} \tau+P_{2} \cos 2 r_{1} \tau, \quad \varphi_{2}=Q_{0}+Q_{1}^{\prime} \cos r_{1} \tau+Q_{2} \cos 2 r_{1} \tau \\
\psi_{2}=R_{0}+R_{1}^{\prime} \cos r_{1} \tau+R_{2} \cos 2 r_{1} \tau \tag{2.14}
\end{gather*}
$$

The constants $P_{1}^{\prime}, Q_{1}^{\prime}$ and $H_{1}^{\prime}$ as well as $\delta_{2}$ are determined from the
 because of their bulkinese.

It is interéwim, Lu note that the coefficients of the $\lambda$ term. in (2.4) represent particular sulutions of the corresponding systems.

The periudic solutions of the system (1.7) art expressed approximately as

$$
\begin{gather*}
\xi=\lambda \cos r_{1} \tau+\lambda^{2}\left(P_{0}+P_{1}^{\prime} \cos r_{1} \tau+P_{2} \cos 2 r_{1} \tau\right)+\ldots \\
\varphi=\lambda k_{1} \cos r_{1} \tau+\lambda^{2}\left(Q_{0}+Q_{1}{ }^{\prime} \cos r_{1} \lambda+Q_{2} \cos 2 r_{1} \tau\right)+\ldots  \tag{2.15}\\
\psi=\lambda k_{2} \cos r_{1} \tau+\lambda^{2}\left(R_{0}+R_{1}^{\prime} \cos r_{1} \tau+R_{2} \cos 2 r_{1} \tau\right)+\ldots
\end{gather*}
$$

and the period of the motion is

$$
T_{1}=\frac{2 \pi}{r_{1}}\left(1+\delta_{2} \lambda^{2}+\ldots\right)
$$

The tension force $K$ in the string of the pendulum i. of the form

$$
K=c(\rho-l)
$$

or, taking into account (2.15)

$$
K=c\left[b-l+\lambda \cos r_{1} \tau+\lambda^{2}\left(P_{0}+P_{1}^{\prime} \cos r_{1} \tau+P_{2} \cos 2 r_{1} \tau\right)+\ldots\right]
$$

Similarly, it can be established that periodic solutions al:n curserpond to the roots $\pm i r_{a}$ and $\pm i r_{s}$

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